## Stellar Astrophysics, Fall 2024 **PROBLEM SET V**

Deadline: 5PM OF THURSDAY, NOVEMBER 21, 2024

1. Outer envelopes of white dwarfs (15%). In class, we showed that the pressure and temperature in the outer envelope of a white dwarf are related by

$$\frac{P^2}{2} = \left[\frac{16\pi}{3}\frac{ac\,Gk}{\kappa_0\,\overline{m}}\frac{M}{L}\right]\frac{T^{8.5}}{8.5}.$$

(a) (10%) Show that the radiative temperature gradient in the outer envelope is given by

$$\frac{\mathrm{d}T}{\mathrm{d}r} = -\frac{GM\overline{m}}{4.25\,r^2\,k}$$

- (b) (5%) Consider a white dwarf with mass  $M = 0.4M_{\odot}$  and radius  $R = 0.01R_{\odot}$  with an internal temperature of 10<sup>7</sup> K. Estimate the thickness of its outer envelope.
- 2. Cooling time of white dwarfs (10%). Show that the time for a carbon white dwarf of mass M to cool from a high internal temperature to a much lower internal temperature,  $T_I$ , is approximately

$$t = \frac{3}{5} \frac{kT_I}{L} \frac{M}{12m_{\rm H}},$$

where L is the luminosity corresponding to  $T_I$ .